SOAR Quick Observing Guide



Blue Camera	Focus	Comp_Arcs	Q-Flats and Bias	GACAM	GW Observing	Other
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- Setup VPN connections for Goodman, NOAO (IRAF) and GACAM
- NOTE: Minimum of 2 preferably 3 monitors required.
- Start up SOAR vpnc in a terminal
 - vpnc /etc/vpnc/vpnc2soar.conf
 - Goodman: vncviewer -Shared 139.229.15.132 &
 - IRAF: vncviewer -Shared 139.229.15.137:4 &
 - GACAM: vncviewer -Shared 139.229.15.168:1 &
- Prepare Goodman Spectrograph Controls
 - Click on GSP_Main-Shortcut Icon to open up Goodman GUI
 - · Click on the arrow to run the Goodman GUI.
 - Click on Main
 —> Logon
 —> Then select the partner (brazil, noao, unc, msu, etc) and enter the
 password that was provided to you by the Support Scientist
 - Click on User \longrightarrow Home Systems \longrightarrow Select All \longrightarrow Then Home Selected
 - WARNING: Before you home systems, make sure the Goodman electronics has been powered on by the Telescope Operator (TO), and that the rotator angle is at 0 deg. Check with the TO.
 - Make sure the Flexure Compensation is ON
 - In the Target Control Box of the Collimator Focus enter 1000 → Then click on "Set". Wait until the collimator reaches the position 1000 in the actual indicator box.
 - Click on the CCD Readout Speed pull-up menu and select the desired readout mode.
 - Click on the CCD ROI Mode pull-up menu and select the region of interest (ROI).
 - Click and select the Slit.
 - Click and select the Grating.
 - Click and select the Wavelength Mode.

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- Set up Focus Sweep to Determine Best Focus
 - Set Stepsize $\longrightarrow 100$
 - Set # Exposure $\longrightarrow 21$
 - Set Mask $\longrightarrow 0.45''$ Slit
 - Set Readout speed $\longrightarrow 400 \text{ kKz}$
 - Set Exposure Time $\longrightarrow 1$ sec
 - Set Custom CCD ROI Mode
 - Set Parallel Origin $\longrightarrow 1900$
 - Set Parallel Length $\longrightarrow 400$
- Measure results to get best focus
 - specfocus *focus.fits

Note: Before setting focus set value to -2000

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Comparison Lamps

- Set up Instrument Parameters for your science program
 - Set mask \longrightarrow e.g. 1.00" Slit
 - Set Readout speed \longrightarrow e.g. 200 kKz, ATTN 2
 - Set CCD ROI Mode \longrightarrow e.g. Spectroscopy 2x2
 - Set Grating \longrightarrow e.g. SYZY_400
 - Set Correct Mode \longrightarrow e.g. 400WD
- Take Comparison Lamps
 - Turn ON Lamps e.g. HgArNe
 - · Let the Telescope Astronomer know you want to take a Comparison Arc
 - Take 1 image Verify Good then take 5 more
 - Let the Telescope Astronomer know you have finished with the Comp Arcs
 - Turn OFF Lamps

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Take Quartz Flats

- Set Intensity to 100 %
- Turn Quartz Lamp ON
- Set Exp. Time: 1.0 sec
- Verify good image in IRAF before continuing
 - display quartz.fits 1
 - implot quartz.fits 480
- Take 50 more exposures
- Take Bias Images
 - Remember to turn Quartz Lamp OFF
 - Update File Name Base to bias
 - Take 50 exposures

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Goodman Acquisition	Camera (GACA)	M) Calibration				

NOTE: Verify the bottom right of GUI says ARM IN or ARM OUT

\rightarrow If not restart GACAM GUI.

- Click the Withdraw Mask button to remove the mask from the optical path.
- Within GACAM GUI place the ARM IN
- Click on Control \longrightarrow Offsets to Open Offset Window
- Click on Exposure \longrightarrow Exposure to Open Exposure Window
- Click on Scale \longrightarrow User MinMax to display Scaling window
- Right click Check Show Cross-Hair UnCheck Lock Cross-Hair
- Drag to move cross hair to center it vertically over center of slit
- Right click Check Lock Cross-Hair
- Center Target in Cross-Hair
 - Place ARM IN
 - Take OUT Mask
 - Center Object in cross hair
 - · Click on center of object.
 - Click COMPUTE in Offset Window
 - Click APPLY in Offset Window
 - Repeat last 3 steps until Object in center of Cross-Hair
 - Place Mask IN
 - If target is visible in slit then Take ARM OUT
 - OBSERVE !!

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Instrument Configuration

- Grating $\longrightarrow 400 \text{ 1/mm}$
- Order $\longrightarrow 1st$
- Slit $\longrightarrow 1.0''$
- Atmosphere disp. corr.: \longrightarrow Yes
- GW Candidate Identifications
 - Standard Star \longrightarrow **YES**
 - Exposure Time
 - $i \le 19 \longrightarrow 900 \text{ sec}$
 - $i \sim 20 \longrightarrow 1200 \text{ sec}$
 - $i \sim 21 \longrightarrow 1800 \text{ sec}$
 - # of Exposures
 - For initial identification --> 1 exposure per Candidate
 - Only if identified as a kilonova → 2 additional exposures

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Left: A simplified flow-chart for a single realization of a simulated SOAR follow-up of a single GW event, where Ncand is the total number of candidates from an imaging search & discovery program. For the simulations here, Ncand is either 8 or 9, but averages overall to 8.79.



Above: Examples of SOAR spectra from an as- sortment of SNe: A=2010as (Ib/c), B=2015L (I-p), C=2010ae (Ia-p), D=2016bro (Ia), E=2018nw (II), F=2018po (Ia), G=2017ijn (IIn), H=2017hvt (Ia).







Right: Results of the simulation (using 100,000 realizations): histogram of the total number of SOAR ToO interrupts for a single LVO 03 event. (Note that the number of interrupts does not scale exactly as the total duration of interrupt time, since the number of hours per interrupt will vary between the "search & discovery" phase and the follow-up phase of the observations for a given kilonova event.) Left: Results of the simulation (using 100,000 realizations): histogram of the total durations of SOAR ToO interrupt time [in hours] for a single LVO O3 event.



